# The FT-IR Spectrum of HC $^{15}$ NO: The $\nu_1$ , $\nu_2$ , $2\nu_3$ , and $\nu_2 + \nu_3$ Band Systems

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The infrared spectrum of HC<sup>15</sup>NO, an isotopically substituted species of fulminic acid, has been measured in the range 1900–3600 cm<sup>-1</sup> at a resolution of 0.003 cm<sup>-1</sup> with a Bruker IFS 120 HR interferometer. More than 100 subbands have been assigned. Power series coefficients for these transitions are given. A Coriolis resonance between the levels 01002 (l = 0e) and 01010 (l = 1e) allows normally "forbidden" transitions to occur, some of which were observed and assigned. We correlate transition intensities and energies of the resonance system. Variations in the manifold of  $n\nu_5$  states with excitation of other modes are compared. © 1993 Academic Press, Inc.

## INTRODUCTION

The vibrational and rotational energy level structure of highly excited vibrational states in polyatomic molecules is currently of experimental and theoretical significance. The fulminic acid molecule, HCNO, is of interest because of its quasilinearity. It is important to obtain experimental data on its isotopomers in order to study the effects of isotopic substitution on the quasilinear HCN bending mode and also on the other bending degree of freedom and their interactions with the stretching modes. As there is now a relatively efficient synthesis (1), the investigation of the rovibrational spectra of H<sup>13</sup>CNO, HC<sup>15</sup>NO, and H<sup>13</sup>C<sup>15</sup>NO is in progress. The rovibrational spectra in the far infrared have now been measured for HCNO and each of the above-mentioned isotopomers (2–6). Because of the low-lying quasilinear bending mode, the molecule offers a rich vibration–rotation overtone spectrum even below 4000 cm<sup>-1</sup>. We are pursuing the observation and analysis of the spectrum to provide data for analysis of the anharmonic potential energy surface and for molecular dynamics studies of the full four-atomic molecule problem. One example is the analysis of the parent species by Iachello *et al.* (7) using the vibron model.

HC<sup>15</sup>NO has five normal modes, three stretching vibrations of  $\Sigma^+$  symmetry, and two bending modes of II symmetry:  $\nu_1$  is the CH stretch at 3333.032 cm<sup>-1</sup>,  $\nu_2$  is the CN stretch at 2152.014 cm<sup>-1</sup>, and  $\nu_3$  is the NO stretch at 1251.218 cm<sup>-1</sup>. The degenerate CNO bending mode  $\nu_4$  lies at 526.562 cm<sup>-1</sup>, and the degenerate quasilinear HCN bending mode  $\nu_5$  at 223.417 cm<sup>-1</sup>. The present results, covering the range 1900–3600 cm<sup>-1</sup>, thus include the band systems  $\nu_1$ ,  $\nu_2$ ,  $2\nu_3$ , and  $\nu_2 + \nu_3$ .

For the present analysis of the data, we have used the formalism appropriate for a linear molecule (3, 8).

## **EXPERIMENTAL DETAILS**

The  $^{15}$ N substitution was introduced in the last step of the synthesis of 3-phenyl-4-oxoiminoisoxazol-s-(4H)-one (I), which was then pyrolyzed as in Ref. (3) to obtain

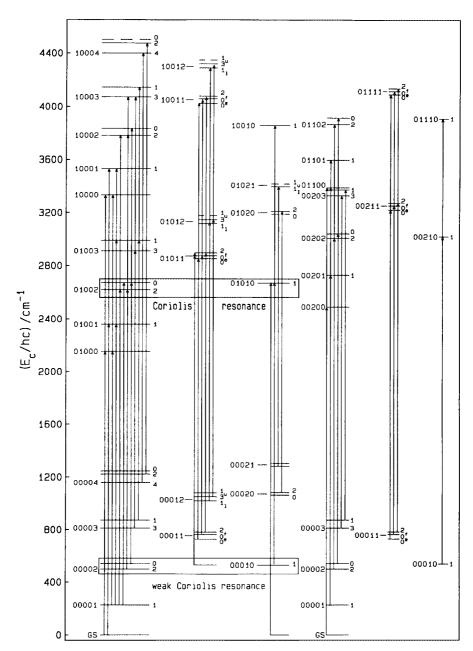


FIG. 1. Vibrational term diagram of HC<sup>15</sup>NO. The box near 2670 cm<sup>-1</sup> indicates the strongest resonance.

 $\rm HC^{15}NO$ . The sample contained approximately 10% unsubstituted HCNO and small amounts of  $\rm H^{15}NCO$  and  $\rm CO_2$ .

The experimental aspects are analogous to those described in Ref. (3). The cell length was 3 m, and the sample was swept into the cell and buffered with argon. The sample handling, as well as Fourier transformation, selection, summing, and obtaining of the transmittance spectra of [15N] fulminic acid were as in Ref. (3). The measure-

TABLE I

Band Centers and Lower States Power Series Constants of HC<sup>15</sup>NO in cm<sup>-1</sup>, Taken from (6)

	$ ilde{ u}_c$	$B_{ps}$	$D_{ps}$	$H_{ps}$	Lps
state			$\times 10^{7}$	$\times 10^{12}$	$\times 10^{15}$
00000	-	0.382 536 85	1.41861	-	-
000011e	223.416 585	0.38295878	1.4014	-0.024	-
$00001^{1f}$	223.416 680	0.38411765	1.45557	-	-
$00002^{2e}$	497.374 040	0.38417535	1.6682	0.022	-
$00002^{2f}$	497.374 040	0.38417458	1.43835	-0.088	-
$00002^{0e}$	540.160 581	0.38403126	1.21075	-0.032	-0.0083
0001 <sup>1e</sup> 0	526.561 639	0.38266340	1.4191	0.407	-0.058
$0001^{1f}0$	526.561 639	0.38341070	1.4367	0.156	-
$000(11)^{0e}$	747.111 034	0.3840564	4.4670	1.1	18.5
$000(11)^{0f}$	748.897 029 3	0.38398967	1.8406	-1.130	0.881
$000(11)^{2e}$	750.111 380	0.3838826	-1.952	-47.6	-
$000(11)^{2f}$	750.111 380	0.38389286	1.0499	3.38	-1.790
$00003^{3e}$	811.002 450	0.384 848 33	1.5583	-0.151	-
$00003^{3f}$	811.002 450	0.38484844	1.5591	0.230	-
$00003^{1e}$	869.528 856	0.38370024	1.3085	0.183	-
$00003^{1f}$	869.528 856	0.38551325	1.3680	-0.188	-
00004 <sup>0e</sup>	1244.346 870	0.384 632 9	3.129	-92.0	-98.3
$00004^{2e}$	1221.742 292	0.3851263	2.313	1.16	-
$00004^{2f}$	1221.742292	0.3851204	1.3173	-	-
$00004^{4e,f}$	1157.813 319	0.3852572	1.5553	-	-
$00100^{0e}$	1251.2175	0.381294	2.47	614.0	-114.7
$000(12)^3$	1024.860 39	0.384 465 4	1.268	-	-
$000(12)^{1_le}$	1020.370 53	0.384572	1.951		-
$000(12)^{1_if}$	1020.36988	0.384 771	1.877	-	

ments were carried out in three regions, in each of which the resolution was chosen so that the full width at half height (FWHH) of the instrumental line shape was about one-half of the Doppler width. Thus, the resolution, given as the inverse of the maximum optical path difference, was  $0.0033~\rm cm^{-1}$  in the region of the optical band-pass  $1850-2350~\rm cm^{-1}$  and  $0.0044~\rm cm^{-1}$  in the band-pass  $2300-3000~\rm cm^{-1}$ . The latter two regions were calibrated with N<sub>2</sub>O (10). The uppermost region,  $3100-3600~\rm cm^{-1}$ , was measured with a resolution of  $0.0055~\rm cm^{-1}$  and calibrated with HCN (11). There is a gap, in which no significant absorption is observed, between  $3000~\rm and~3100~\rm cm^{-1}$ , which was not measured.

## ASSIGNMENT AND ANALYSIS OF THE SPECTRUM

The vibrational transitions observed for HC<sup>15</sup>NO are summarized in Fig. 1 which displays the vibrational term diagram. The assignments were made initially in analogy with those of the parent species of HCNO (12–14) and confirmed using lower-state combination differences or by comparing the lower-state rotational parameters with those obtained in Refs. (5, 6). The observation of many vibrational states two or three times over, in transitions from the ground state and from various excited levels, confirms many of the assignments. The constants of most of the lower vibrational states of HC<sup>15</sup>NO are taken from Ref. (6); they are all given in Table I. Transitions from the far infrared, used to determine the constants for the 00012 levels, were taken from the spectra of Refs. (5, 6) and are included in Appendix E<sup>1</sup> of this paper.

Appendices comprising 47 tables have been kept on deposit in the Editorial office of the *Journal of Molecular Spectroscopy*. Copies may be obtained from the authors upon request.

The assignment process is done efficiently with an interactive Loomis-Wood assignment program (15) and a similarly interactive program to check combination differences. The initial data reduction procedure for the subbands, characterized by a given set of vibrational quantum numbers  $(v_1v_2v_3v_4v_5)$ , a given set of vibrational angular quantum numbers  $l_4$ ,  $l_5$  and  $k = l_4 + l_5$ , and a given parity e or f, is to fit each subband separately. A given state is indicated by  $v_1v_2v_3v_4^lv_5^l$  or  $v_1v_2v_3$  ( $v_4v_5$ )<sup>k</sup>. Two different combinations of  $l_4$  and  $l_5$  giving the same k are distinguished by a subscript l for the lower and u for the upper of the two levels. The data reduction is done by a program which calculates band centers and effective rotational and centrifugal distortion constants, that is, coefficients in a power series (ps) expansion of the rovibrational term value in J(J+1),

$$E(J) = G_c + B_{ps}J(J+1) + D_{ps}J^2(J+1)^2 + H_{ps}J^3(J+1)^3 + L_{ps}J^4(J+1)^4, (1)$$

for both the upper and the lower states. The band center  $\tilde{\nu}_c = (G'_c - G''_c)$  is defined as the difference between term values which are thus different from the spectroscopic

TABLE II

Band Centers and Upper State Power Series Constants of Subbands in the  $\nu_2$  Band System of HC<sup>15</sup>NO in cm<sup>-1</sup>

subband	$\tilde{ u}_c$	$B'_{ps}$	D <sub>ps</sub> ×10 <sup>6</sup>	$H'_{ps} \times 10^{12}$	$L'_{ps} \times 10^{15}$	σ ×10 <sup>4</sup>	lines used
01000°e ← 00000°e	2152.00881 (20)	0.380 422 5 (22)	0.5073 (75)	294.6 (96)	-89.2 (40)	4.182	120
01000° ← 00001¹/	1928.59053 (44)	0.380 421 7 (49)	0.522 (18)	330.1 (240)	-110.8 (110)	2.893	Q 33
01000° ← 000011e	1928.59032 (34)	0.380 417 0 (34)	0.478 (10)	248.2 (110)	-67.6 (40)	6.154	71
$01001^{1f} \leftarrow 00000^{0e}$	2356.591497 (12)		0.145 13 (30)	- 10.2 (110)	-	0.1832	
010011e ← 000000e		0.380 550 964 (92)	0.139 739 (87)	•	-	0.7388	97
010011e ← 000011e	2133.172 19 (27)	0.380 550 470 (83)	0.139 393 (53)	-0.0419 (92)	-	0.9287	109
$01001^{1/} \leftarrow 00001^{1/}$	2133.172 16 (28)	0.381766945 (80)	0.144793 (47)	0.0202 (74)	-	1.0614	120
$01002^{2e} \leftarrow 00001^{17}$	2393.302.866 (99)		0.1584 (67)	0.64 (29)			
$01002^{2e} \leftarrow 00001^{1e}$	2393.302953 (32)		0.157 27 (13)	-	-	1.168	69
$01002^{2f} \leftarrow 00001^{1e}$	2393.302902 (47)		0.13963 (10)		_	0.9664	
$01002^{2f} \leftarrow 00001^{1f}$	2393.303.007 (24)		0.139 260 (96)	-	_	0.8966	70
01002 <sup>2e</sup> ← 00002 <sup>2e</sup>	2119.342 674 (60)		0.137 94 (29)	-	-	2.373	101
$01002^{2f} \leftarrow 00002^{2f}$	2119.34283 (10)	0.38185761 (68)	0.1573 (12)	63.8 (67)	_	3.433	105
010020e 000000e	2669.6967 (11)	0.382364 (10)	1.080 (30)	656.6 (320)	-166.9 (110)		88
$01002^{0e} \leftarrow 00001^{1f}$	2446.2754 (14)	0.382426 (15)	1.320 (52)	974.5 (640)	-297.2 (250)		Q 35
010020e 000011e	2446.27836 (90)	0.382395 (11)	1.226 (37)	869.5 (480)	-256.6 (190)	7.076	39
$01002^{0e} - 00002^{0e}$	2129.5313 (15)	0.382308 (13)	0.821 (28)	349.5 (210)	-62.1 (48)	55.94	87
$01002^{0e} \leftarrow 0001^{1e}0$	2143.1463 (12)	0.382 265 1 (95)	0.864 (25)	485.5 (250)	-125.5 (890)	4.621	36
$01003^{1e} \leftarrow 00003^{1e}$	2119.348134 (98)		0.1894 (19)	10.0 (16)	-1.02 (44)	2.202	75
$01003^{1f} \leftarrow 00003^{1f}$	2119.347.86 (11)	0.3834396(9)	0.1574 (25)	6.1 (25)	-1.53 (81)	1.132	56
010031c - 000011c	2765.461 234 (76)		0.1769 (43)	- (/	-	1.813	46
$01003^{1f} \leftarrow 00001^{1f}$	2765.460652 (81)		0.15072 (45)	-	_	2.180	56
$01003^3 \leftarrow 00003^3$	2109.009712 (94)		0.1455 (16)	-8.7 (13)	2.06 (31)	2.126	79
$01004^4 \leftarrow 00004^4$	2102.3138 (17)	0.383 208 4 (12)	0.1577 (25)	17.8 (23)	(/	1.701	35
$0101^{1e}0 \leftarrow 0001^{1e}0$	2142.6435 (14)	0.379672 (12)	-0.606 (28)	-399.2 (230)	77.4 (60)	48.01	91
$0101^{1e}0 \leftarrow 00002^{0e}$	2129.0425 (13)	0.379621 (13)	-0.894 (40)	-802.8 (480)	240.9 (190)		43
01011e0 ← 000011e	2445.7933 (12)	0.379 555 (13)	-1.077 (44)	-995.3 (570)	307.2 (240)		60
$0101^{1f}0 \leftarrow 0001^{1f}0$	2142.662013 (27)		0.16493 (61)	7.77 (54)	-1.14 (15)	0.8330	90
$0101^{1f}0 \leftarrow 00001^{1f}$	2445.80721(12)	0.381 106 9 (12)	0.1737 (40)	20.1 (47)	-5.1 (18)	1.918	67
$0101^{1f}0 \leftarrow 00000^{0e}$	2669.22404 (10)	0.381 108 82 (86)	0.1792 (20)	15.9 (13)	- '	1.058	Q 23
01011 <sup>0e</sup> ← 00011 <sup>0e</sup>	2123.912.856 (70)		0.4296 (18)	-1.92 (17)	16.6 (53)	1.625	70
$01011^{0f} \leftarrow 00011^{0f}$	2123.553404 (44)		0.168 25 (58)	-3.47 (36)	0.95 (7)	1.458	75
$01011^{0f} \leftarrow 0001^{1e}0$	2345.888 68 (37	0.3816216(29)	0.1792 (50)	-` ′	- ' '	2.878	Q 14
$01011^{2e} \leftarrow 00011^{2e}$	2124.24408 (12)	0.3815105(11)	-0.1737 (21)	-33.6 (14)	-2.51 (16)	3.407	58
$01011^{2e} \leftarrow 0001^{1f}0$	2347.79401 (55)	0.3815103 (43)	-0.1713 (83)	-35.6 (44)	•	8.222	Q 28
$01011^{2f} \leftarrow 00011^{2f}$	2124.243.85 (19)	0.381 521 6 (12)	0.1153 (19)	7.68 (11)	-2.44 (21)	3.728	81
$01011^{2f} \leftarrow 0001^{1e}0$	2347.794 38 (60)	0.3815219 (36)	0.1123 (44)	-	-	7.671	Q 19
$01012^3 \leftarrow 00012^3$	2111.248652 (63)		0.11749 (91)	6.09 (50)		1.132	54
$01012^{1_i\epsilon} \leftarrow 00012^{1_i\epsilon}$	2110.95562 (15)	0.38245227 (68)	0.16707 (68)	-	_	4.985	47
$01012^{1_if} \leftarrow 00012^{1_if}$	2110.95683 (11)	0.38258050 (64)	0.164 24 (67)	_	_	3.700	52
			= = = = = (01)			3	

subband	$\tilde{\nu}_c$	$B_{ps}^{\prime}$	$D_{ps}$	$H_{ps}^{'}$	$L_{ps}$	σ	lines
		•	$\times 10^6$	$\times 10^{12}$	$\times 10^{15}$	$\times 10^4$	used
00200 <sup>0e</sup> ← 00000 <sup>0e</sup>	2491.54205 (21)	0.379 186 3 (13)	0.1819 (19)	3.37 (97)	0.71 (16)	8.666	107
002011e ← 000011e	2502.280882 (70)	0.379 599 28 (62)	0.1461 (16)	3.2 (15)	-0.91 (45)	1.754	97
$00201^{1f} \leftarrow 00001^{1f}$	2502.280878 (27)	0.38072472 (17)	0.148 35 (30)	0.18 (14)	-	0.827	97
$00202^{0e} \leftarrow 00002^{0e}$	2501.093754 (59)	0.3807184 (52)	0.1210 (15)	6.3 (15)	-1.20(49)	0.841	74
$00202^{2e} \leftarrow 00002^{2e}$	2509.34099 (40)	0.3808448 (68)	0.130 (33)	95.7 (480)	-	2.591	29
002031e ← 000031e	2502.4774 (11)	0.380 289 3 (90)	-0.33 (19)	-	-	4.081	16
$00203^{1f} \leftarrow 00003^{1f}$	2502.47267 (68)	0.382 146 3 (45)	0.034(7)	-	-	5.778	24
$00203^{3e} \leftarrow 00003^{3e}$	2514.50405 (90)	0.381 654 9 (98)	0.276 (31)	214.7 (290)	-	3.775	31
$0021^{1e}0 \leftarrow 0001^{1e}0$	2492.045780 (41)	0.3794464 (39)	0.2447 (11)	52.6 (12)	-9.12 (38)	0.775	79
$0021^{1f}0 \leftarrow 0001^{1f}0$	2492.045783 (59)	0.380 207 9 (57)	0.2473 (16)	49.2 (17)	-9.24(59)	1.113	73
$00211^{0e} \leftarrow 00011^{0e}$	2503.41282 (28)	0.3806655(25)	0.534 (66)	67.8 (50)	-	2.830	42
$00211^{0f} \leftarrow 00011^{0f}$	2502.94717 (15)	0.3806075(17)	0.1575 (61)	-14.6 (79)	5.7(33)	1.776	48
$00211^{2e} \leftarrow 00011^{2e}$	2503.224284 (95)	0.380 521 5 (15)	-0.2059 (41)	-50.2 (27)		4.625	50
$00211^{2f} \leftarrow 00011^{2f}$	2503.224343 (82)	0.380 524 9 (19)	0.1287 (51)	4.5 (29)	-	4.761	55
$00212^{1_1e} \leftarrow 00001^{1e}$	3307.81481 (16)	0.381 245 56 (92)	0.1909 (15)	42.76 (87)	-4.48 (16)	4.201	81
$00212^{1_1e} \leftarrow 00000^{0e}$	3530.85049 (15)	0.381 232 26 (97)	0.1572 (17)	17.24 (85)	- ` `	2.044	45
$00212^{1_if} \leftarrow 00001^{1f}$	3307.81691 (38)	0.381 807 5 (27)	-0.0036 (54)	160.6 (39)	-47.42 (92)	9.467	86
$00212^{\mathbf{I}_{i}f} \leftarrow 00000^{0\epsilon}$	3531.22543 (38)	0.3817381 (36)	-0.1294 (91)	59.2 (65)	-	3.877	Q 24

TABLE IV Band Centers and Upper State Power Series Constants of Subbands in the  $\nu_1$  Band System of HC  $^{15}$ NO in cm $^{-1}$ 

subband	$\tilde{\nu}_{\rm c}$	$B'_{ps}$	$D'_{ps}$	$H_{ps}$	$L_{ps}^{'}$	σ	lines
		•	$\times 10^6$	$\times 10^{12}$	$\times 10^{15}$	$\times 10^{4}$	used
10000 ← 00000	3333.030617 (14)	0.381 300 842 (68)	0.158 437 (82)	-0.908 (26)	-	0.576	95
$10000^{0e} \leftarrow 00001^{1e}$	3109.613881 (64)	0.381 300 59 (36)	0.158 91 (44)	- '	-	1.410	52
$10000^{0e} \leftarrow 00001^{1f}$	3109.613901 (61)	0.381 299 45 (45)	0.157 76 (94)	-1.21(56)	-	0.785	Q 35
$10001^{1e} \leftarrow 00000^{0e}$	3529.10420 (11)	0.381 696 2 (13)	0.1871 (41)	-33.4 (49)	9.3 (19)	2.325	73
$10001^{1e} \leftarrow 00001^{1e}$	3305.687581 (69)	0.381 693 96 (79)	0.1843 (27)	-33.7(34)	8.4 (14)	1.320	65
$10001^{1f} \leftarrow 00001^{1f}$	3305.687727 (74)	0.3827986 (10)	0.2954 (44)	-166.1 (71)	49.5 (36)	1.055	55
$10002^{2e} \leftarrow 00001^{1e}$	3557.245 528 (64)	0.383 237 85 (35)	0.166 78 (40)	-	-	1.550	39
$10002^{2e} \leftarrow 00001^{1f}$	3557.245 558 (76)	0.383 236 38 (33)	0.166 10 (27)	-	-	1.634	Q 31
$10002^{2f} \leftarrow 00001^{1f}$	3557.245 552 (91)	0.383 236 24 (51)	0.146 36 (57)	-	-	2.392	44
$10002^{2f} \leftarrow 00001^{1e}$	3557.245 21 (12)	0.383 238 43 (47)	0.147 55 (35)	-	-	1.621	Q 24
$10002^{2e} \leftarrow 00002^{2e}$	3283.28797 (14)	0.383 235 40 (63)	0.165 09 (60)	-	-	4.347	77
$10002^{2f} \leftarrow 00002^{2f}$	3283.28778 (11)	0.383 236 89 (51)	0.146 18 (46)	-	-	4.023	76
$10002^{0e} \leftarrow 00002^{0e}$	3292.996603 (24)	0.382 897 81 (12)	0.12869 (12)	-	-	0.771	77
$10003^{3e} \leftarrow 00002^{2f}$	3571.30173 (15)	0.3837498 (14)	0.1857 (29)	-	_	1.919	Q 17
$10003^{3e} \leftarrow 00002^{2e}$	3571 301 29 (23)	0.3837533(30)	0.2078 (82)	-	-	2.429	14
$10003^{3f} \leftarrow 00003^{3f}$	3257.672738 (91)	0.3837527 (77)	0.1950(18)	-6.3(15)	1.56 (41)	2.165	67
$10003^{3f} \leftarrow 00003^{3e}$	3257.67290 (75)	0.3837506 (82)	0.199 (18)		- ' '	10.51	Q 16
$10003^{1e} \leftarrow 00003^{1e}$	3275.327 395 (99)	0.38251460 (50)	0.11947 (46)	-	-	2.397	36
$10003^{1f} \leftarrow 00003^{1f}$	3275.327550 (83)	0.38448130(42)	0.131 12 (40)	-	-	2.246	40
$10004^{4e} \leftarrow 00004^{4e}$	3239.317513 (92)	0.384 295 47 (45)	0.161 35 (48)	-	-	1.549	50
$10004^{2e} \leftarrow 00004^{2e}$	3247.31761 (38)	0.383 916 6 (28)	0.2760 (44)	-	-	7.499	33
$10004^{2f} \leftarrow 00004^{2f}$	3247.31780 (41)	0.383 905 1 (31)	0.1526 (49)	-	-	7.937	32
$1001^{1e}0 \leftarrow 0001^{1e}0$	3329.41716 (33)	0.381 386 1 (31)	0.1610 (76)	-15.2 (54)		4.925	42
$1001^{1/}0 \leftarrow 0001^{1/}0$	3329.41482 (31)	0.3821024(28)	0.2321 (69)	42.6 (48)	-	4.477	45
$10011^{0e} \leftarrow 00011^{0e}$	3302.61851 (15)	0.3828733(12)	0.6076(25)	57.9 (14)	-	3.065	45
$10011^{0f} \leftarrow 00011^{0f}$	3302.22253 (12)	0.38272767 (52)	0.237 79 (43)	- '	-	4.070	50
$10011^{2e} \leftarrow 00011^{2e}$	3302.347 19 (41)	0.3824924(47)	-0.249 (15)	-101.0 (130)	-	6.835	43
$10011^{2f} \leftarrow 00011^{2f}$	3302.347.29 (26)	0.382 502 9 (30)	0.1244 (84)	-65.9(61)	-	4.773	38
$10012^3 \leftarrow 00012^3$	3279.785 19 (24)	0.383 526 8 (23)	0.1035 (51)	-6.5(30)	-	5.816	55
$10012^{1_1e} \leftarrow 00012^{1_1e}$	3280.40270 (24)	0.383 685 3 (22)	0.2287(52)	20.7 (41)	-4.7(10)	5.800	52
$10012^{1,f} \leftarrow 00012^{1,f}$	3280.405 46 (31)	0.3837738 (14)	0.1936 (16)			3.096	50

TABLE V
Band Centers and Upper State Power Series Constants of Subbands in the $\nu_2 + \nu_3$ Band System of HC <sup>15</sup> NO in cm <sup>-1</sup>

subband	$\tilde{\nu}_c$	$B_{ps}^{\prime}$	$D_{p_s}$	$H_{ps}'$ $\times 10^{12}$	$L_{ps}$	σ	lines
		, .	$\times 10^6$	$\times 10^{12}$	$\times 10^{15}$	$\times 10^4$	used
011000e ← 000000e	3387.485 33 (20)	0.378 835 5 (13)	0.0756 (23)	-20.4 (14)	8.87 (26)	7.303	89
$01100^{0s} \leftarrow 00001^{1f}$	3164.06824 (39)	0.378 847 3 (36)	0.1091 (94)	10.5 (73)	- '	2.496	Q 17
$01101^{1e} \leftarrow 00001^{1e}$	3374.02300 (36)	0.379 053 8 (17)	0.1235 (17)	-	-	9.456	47
$01101^{1f} \leftarrow 00001^{1f}$	3374.02299 (24)	0.380 262 4 (21)	0.1076 (47)	-14.8 (29)	-	3.937	39
$01102^{2e} \leftarrow 00002^{2e}$	3366.6398 (19)	0.380 247 (21)	-0.435 (72)	-526. (88)	166. (35)	26.75	57
$01102^{0e} \leftarrow 00002^{0e}$	3372.67581 (22)	0.380 243 5 (31)	0.047 (13)	-95.2 (190)	76.4 (94)	3.405	45
01111e0 - 00011e0	3379.03953 (12)	0.379 033 9 (12)	0.0925 (35)	-6.9 (38)	7.7 (13)	2.117	64
$0111^{1/9}0 \leftarrow 0001^{1/9}0$			0.092 98 (64)	-	-	4.382	59
0111110e - 000110e	3365.9225 (85)	0.380 252 4 (84)	0.628 (23)	145.4 (190)	-	11.46	32
$01111^{of} \leftarrow 00011^{of}$	3365.37537 (30)	0.380 133 4 (15)	0.1473 (16)	-	-	5.260	32
$01111^{2e} \leftarrow 00011^{2e}$	3366.1053 (12)	0.380 101 (14)	-0.089 (23)	-	-	19.33	19
$011111^{2f} \leftarrow 00011^{2f}$	3366.1123 (24)	0.380 031 (15)	0.024 (23)	-	-	46.26	23

term value  $G_v$  (8) which is defined by band origins. When the lower state had been identified, the constants for the lower state were held fixed at the values given in Table I while constants of the upper state were determined. The expression in Eq. (1) cannot describe a complete k polyade of an excited state simultaneously. The power series coefficients provide only a nearly "model-free" representation of the term values. They are a check of the vibrational assignment, but are not directly related to a molecular Hamiltonian.

In earlier papers reporting data for fulminic acid and its isotopomers we have fit all the data with an effective Hamiltonian for linear molecules (8). In contrast to the situation in the far IR, this ansatz gives satisfactory results in only a few exceptional cases, for states with term values about  $2000 \text{ cm}^{-1}$ . The perturbations observed here are so numerous, that it was not possible to determine physically significant constants in the Hamiltonian for more than a fraction of the bands reported here. Even where we have identified the resonances, a suitable Hamiltonian would require the simultaneous fit of transitions to several states with various l multiplets. We have for the time being limited ourselves to extracting physical information from the band centers and  $B_{ps}$  values. Even relying on the simple analysis based on Eq. (1), the density of states in the range of energy investigated is already so large that resonances with normally "dark" states distort the bands, leading to deviations up to 100 times the experimental standard deviation  $\sigma$ .

The constants obtained for the upper states involved in the rovibrational transitions of the  $\nu_1$ ,  $\nu_2$ ,  $2\nu_3$ , and  $\nu_2 + \nu_3$  band systems so far assigned are given in Tables II-V. The complete list of line positions of the assigned rovibrational transitions is given in Appendices A-D in a sequence corresponding to Tables II-V. The fit procedure used a relative weighting factor w for every line with  $w = (1 - \text{transmittance})^2$ , thus giving a very low weight to the far wings of the bands. Therefore in some perturbed bands, large deviations are found for lines with high J values.

The combination of the effects of accidental resonances and irregularities resulting from the quasilinearity of the  $\nu_5$  manifold leads to a complex spectrum. Many of the bands show localized rotational perturbations; an example of a global Coriolis perturbation of entire bands is presented below. The spectra of HCNO and its isotopomers also show a pattern of Fermi resonances (6). There is almost no band that is free of resonance effects. The constants in Tables II–V are affected by all the known and

unknown resonances, although we have excluded from the fit lines obviously perturbed by local resonances. In Table VI we list the states identified so far, their term values as defined by Eq. (1), and comments concerning perturbed states. The horizontal lines in Table VI delineate interacting groups of levels. As yet no systematic assignment of the partners in the local resonances has been attempted, although several have been identified.

#### "FORBIDDEN" BANDS

As already noted, there are many resonances in the spectrum of HC  $^{15}$ NO. In Fig. 1 a box around the interacting terms shows a prominent resonance found in the spectrum. A Coriolis resonance between the states  $0101^{16}$ 0 and  $01002^{06}$  gives rise to several normally forbidden bands, or more rigorously, Coriolis-allowed transitions. The centers of the bands arising from the terms in resonance have a separation of 0.5 cm<sup>-1</sup>. The transitions  $0101^{16}$ 0  $\leftarrow 00002^{06}$  and  $01002^{06} \leftarrow 0001^{16}$ 0, which actually acquire intensity due to a Coriolis interaction both in the lower state (small effect) (6) and in the upper state (dominant effect), were found. They are shown in Fig. 2 as diagonal arrows which are omitted in Fig. 1. The subbands involved show especially obvious deviations in the power series fits.

Figure 3 displays in a Loomis-Wood diagram one band of this resonance, the hot band  $0101^{1e,f}0-0001^{1e,f}0$ , in which the *e* component shows the effect of the  $0101^{1e}0/10001^{1e}$ 

TABLE VI

Term Values Determined or Estimated for HC<sup>15</sup>NO in cm<sup>-1</sup>
(States in the Same Box May Be in Resonance)

STATE	$G_{\mathfrak{c}}$	INTERACTIONS	STATE	$G_c$	INTERACTIONS
00015°	2158 ??		10001	3529.104	str.r.
01000	2152.015	gl.r.	$002(12)^{1_I}$	3531.225	str.r.
01001	2356.591		01101	3597.440	l.r.J=12, str.l.r.J=15/16
00200	2491.542	gl.r., l.r. J=47	100022	3780.662	
$01002^2$	2616.720		10002°	3833.158	
01010	2669.224	Cotiolis gl.r.	10010	3855.979	
$01002^{0}$	2669.697	without crossing	01102 <sup>2</sup>	3864.006	gl.r., str.l.r. J=13
00201	2725.698	l.r. J=27	01110	3905.602	
010(11)00	2871.024		01102°	3912.837	l.r. J=31
$010(11)^{0f}$	2872.450		100(11)0e	4049.730	
$010(11)^2$	2874.355	gi.r.	100(11)0/	4051.120	
$01003^3$	2920.012	l.r. J=34	100(11) <sup>2</sup>	4052.458	i.r. J=21
010031	2988.878		100033	4068.675	l.r. J=24
$00202^{2}$	3006.715	str.gl.r., l.r. J= 29	011(11)0e	4113.038	l.r.
00210	3018.608	I.r. J=18	011(11)0f	4114.272	i.r.
00202°	3041.255	l.r. J=42	$011(11)^2$	4116.566	str.l.r. J=13
$010(12)^{1_i}$	3131.326		100031	4144.856	
$010(12)^3$	3136.109		100(12),1	4300.774	
$002(11)^{0e}$	3250.524		$100(12)^3$	4304.645	gl.r.
$002(11)^{0f}$	3251.844		1002 <sup>2</sup> 0	4378.201	
$002(11)^2$	3253.355		10004 <sup>4</sup>	4397.131	ļ
010044	3260.127		10004 <sup>2</sup>	4469.060	Ĺ
$00203^3$	3325.506	gl.r.			
10000	3333.031	weak gl.r.			
$00203^{1}$	3371.999				
$01004^{a}$	3383. ??				
01100	3387.485	str.r., l.r. J=30, 48			

str: strong, gl: global, l: local, r: resonance,

<sup>??:</sup> expected value

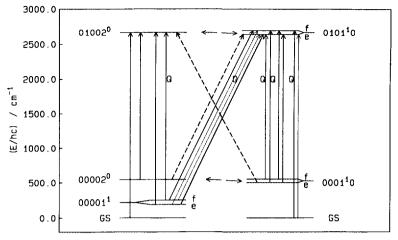


FIG. 2. Transitions connecting the interacting levels 00010/00002 and 01010/01002 in HC <sup>15</sup>NO. Full lines show allowed transitions. The thin lines indicate transitions which are allowed, but not seen. The solid diagonal arrows involving the e levels in the upper state are resonance enhanced, while broken diagonal arrows are Coriolis resonance-induced transitions.

 $01002^{0e}$  interaction. The difference band  $0101^{1e,f}0 \leftarrow 00001^{1e,f}$  must be classified as resonance enhanced for the e component, since even the f component, which is unaffected by either resonance, could be seen. However, the P branch of the e component is a factor of roughly 1.8 stronger and the R branch of the e component is the same factor weaker than the f component, as a result of resonance enhancement/depletion of the transition moment.

A plot of the reduced rovibrational term values for the transitions to the upper states  $0101^{10}$  and  $01002^{0}$  from J = 1 to J = 60, reproduced in Fig. 4, shows that no crossing of the term values of these states takes place. This is in contrast to the pattern

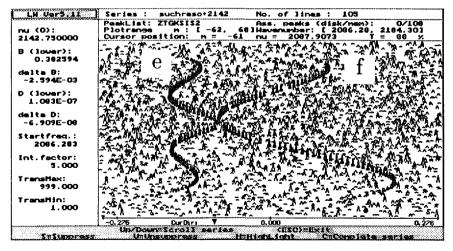


FIG. 3. Coriolis perturbed (e) and unperturbed (f) components of the  $0101^{1e,f}0-0001^{1e,f}0$  subbands in HC<sup>15</sup>NO in a Loomis-Wood program screen dump.

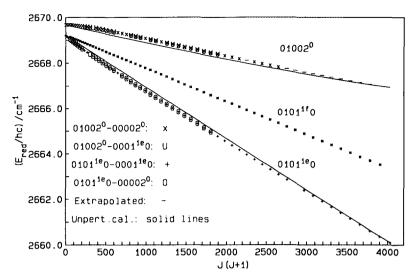


FIG. 4. Reduced term values of the 01010/01002 levels of HC<sup>15</sup>NO. The full lines are unperturbed levels calculated with the Coriolis resonance coefficient set to zero.

of the  $00010^{1e}/00002^{0e}$  resonance in H<sup>13</sup>CNO, where the analogous unperturbed levels cross (4). However, since the divergence of the unperturbed levels is slow, and the interaction matrix element increases with J, it is a global perturbation of both states, for all observed J values. The observed reduced term values were obtained using the Ritz combination principle. The unperturbed calculated curves in Fig. 4 are calculated with the adjusted set of fitted constants, determined from the effective Hamiltonian including the resonance as in Ref. (4) for both upper and lower states, but with the Coriolis coupling constant set to zero.

Figure 5 shows a plot of the absorbtance values of some of the transitions to the interacting levels indicated in Fig. 2. The plot shows an equal absorbtance for both of the allowed hot bands, which is to be expected, since the Bolzmann factor for the two bands is indistinguishable. Similarly, the plot shows an equal absorbtance of the two resonance-allowed transitions, which are rather weak relative to the allowed hot bands: they are probably a factor of 10 weaker than the allowed hot bands. This is consistent with the information revealed by Fig. 4: The shift of the term values is only moderate, so the mixing of the levels leads to only weakly allowed resonance-induced transitions. It is to be noted, especially, that P and R branches are equally intense in the resonance-allowed bands. This indicates (see the discussion in Ref. (4)) that the transition moments for the unperturbed transitions  $01002 \leftarrow 00010$  and  $01010 \leftarrow 00002$  are effectively zero. This again is in contrast to the situation found in Ref. (4), and to that of the difference band  $0101^{1e}0 \leftarrow 00001^{1e}$ , mentioned above.

A further obvious resonance system involves the  $10001^{-1}$  state. Two bands of nearly equal strength appear in the region of both the hot band  $10001^{-1} \leftarrow 00001^{-1}$ , near 3306 cm<sup>-1</sup>, and the combination band  $10001^{-1} \leftarrow 00000$ , near 3531 cm<sup>-1</sup>, involving both e and f components. These four bands define a resonance for the upper state,  $10001^{-1}$ , which does not have a counterpart in the parent species. The tentative assignment of the interacting level is  $002(12)^{1/e^{-f}}$  or  $002(1^{-1}2^{-2})$ . Figure 6 is a plot of the reduced rovibrational term values for the upper state. It can be interpreted as showing crossings of the term values for the f components near  $J \approx 30$  and at higher J (>40) for the

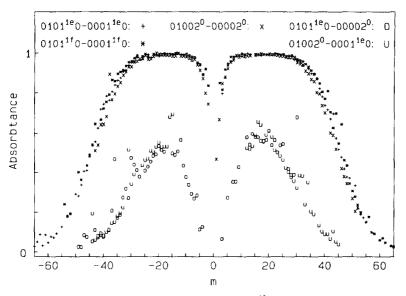


Fig. 5. Absorbtance of the 01010/01002 hot band system in HC <sup>15</sup>NO. The absorbtance values are plotted versus the index m, where m = -J in the P branch and m = J + 1 in the R branch.

e components. The identification indicates the nature of many of the interactions which we have to consider: in particular, high excitation of one or both bending modes.

Numerous further subbands, for which we could not yet find a reliable vibrational assignment, have also been observed. We hope to assign some of them on the basis of improved predictions of "dark" states.

## DISCUSSION

Figure 7 shows the term values of the anharmonic HCN bending mode  $\nu_5$  of HC<sup>15</sup>NO as a function of the excitation of the remaining vibrational modes. The CH stretch  $\nu_1$ 

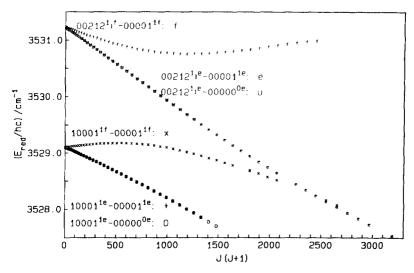


Fig. 6. Reduced term values of the 10001/00212 levels of HC<sup>15</sup>NO.

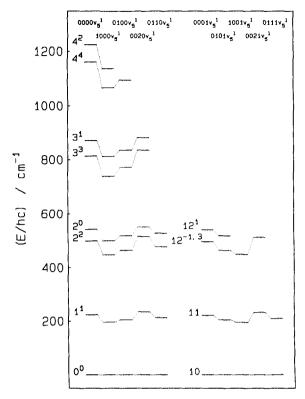


Fig. 7. Dependence of the term value intervals in the  $\nu_5$  manifold on the other modes in HC<sup>15</sup>NO.

has the strongest influence on the HCN bending levels, lowering the vibrational manifold and increasing the splitting, i.e., increasing the quasilinear character of the  $\nu_5$  stack. This indicates an increase of the barrier height in the HCN bending potential upon excitation of the C-H stretching mode. The excitation of the NO stretch overtone  $2\nu_3$  leads on the other hand to a slight upward shift in energy of the HCN bending levels. Thus, excitation of this latter mode decreases the quasilinear character of the molecule. The shifts per quantum of  $\nu_3$  are found to be larger if  $\nu_2$  is excited; in a slightly anharmonic model, this would indicate a nonzero value of  $x_{23}$ . The anharmonic intervals and shifts summarized in Fig. 7 correspond roughly to those found from more limited data for the parent species (9). They indicate why it is difficult to extrapolate upwards to identify resonance partners in this molecule.

Figure 8a shows large and very nonlinear variations of the rotational constant  $B_{ps}$  as a function of the excitation of the bending vibrational mode  $\nu_5$  and of the other modes. The reduced presentation emphasizes the deviation of the B values in HC <sup>15</sup>NO from the linear expansion for the harmonic case

$$B_v = B_e - \sum_i \alpha_i (v_i + d_i/2).$$
 (2)

These deviations parallel the effects shown in Fig. 7. We report in Fig. 8b a similar summary for the l-type doubling constant  $q_5$ , which describes the splitting of the l = 1 levels according to



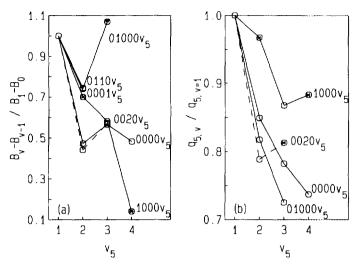


Fig. 8. Dependence of (a) the rotational constants  $B_{ps}$  and (b) the l-type doubling constant  $q_5$  on the bending quantum number  $v_5$  in different vibrational states of HC <sup>15</sup>NO. For l > 0, the average of  $B_{ps}$  for the e and f values is plotted. Asterisks indicate values derived for perturbed levels, which fall out of the general pattern.

$$\Delta E = \frac{1}{2}q_5(v_5+1)J(J+1). \tag{3}$$

For  $v_5 = 2$  and 4, values of  $q_5$  were obtained from fits to the effective Hamiltonian (3, 8), which, as noted above, were otherwise of limited value. As can be seen, this quantity is not nearly constant, as would be expected for a linear molecule, but changes by 20% from the first to the third excited state.

An overview of the spectrum of HC<sup>15</sup>NO is given in Figs. 9 and 10. Many of the bands reported here can be identified in these spectra with the aid of band centers listed in Tables II–V. Some transitions are indicated with an assignment comb. Because of the chemical instability of HC<sup>15</sup>NO the isomerization product H<sup>15</sup>NCO is also present. Also, 10% of the sample is the main species HCNO, while H<sup>13</sup>C<sup>15</sup>NO is present in the natural abundance of <sup>13</sup>C ( $\sim$ 1%), both of which contribute observable bands. CO<sub>2</sub> is also observable.

The intercomparison of the spectra of HCNO and HC <sup>15</sup>NO has aided assignments in both spectra. When more of the resonances have been analyzed, a comprehensive picture of the energy levels and resonances can be filled in for both species.

Since the nitrogen nucleus in fulminic acid lies nearly at the center of gravity, one expects that the species HC  $^{15}$ NO should behave quite similarly to HCNO rotationally, but not necessarily vibrationally, especially for modes involving the nitrogen. The  $\nu_1$  transitions and their combination bands and hot bands in the four species HCNO, HC  $^{15}$ NO, H  $^{13}$ CNO, and H  $^{13}$ C  $^{15}$ NO lie near each other and are very similar. The most striking difference between these bands in HC  $^{15}$ NO and HCNO is the strong resonance pattern of the 10001 state, discussed above, which is brought into resonance by the  $^{15}$ N isotope shift of the 00212 state. The transitions of hot bands of the  $\nu_2$  combination states in which a CN vibration is excited are shifted downward in wavenumber in the  $^{15}$ N species by ca. 44 cm $^{-1}$ . The downward shift of the  $\nu_2$  levels and the shift in  $\nu_4$  relative to  $2\nu_5$  lead to an accidental separation of less than 1 cm $^{-1}$  for the resonance system  $(02)^0$  and  $(10)^{10}$  in combination with  $\nu_2$ . In the spectra of the

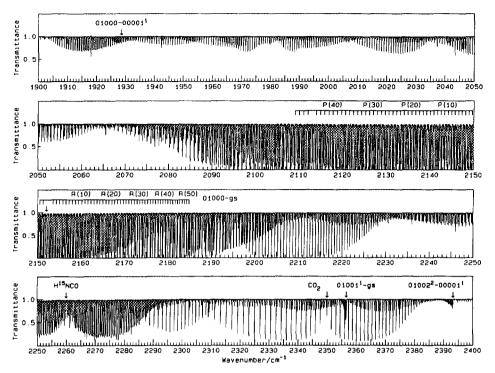


FIG. 9. Spectrum of HC15NO in the range 1900-2400 cm<sup>-1</sup>.

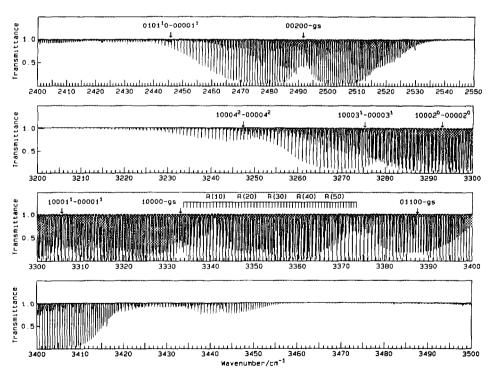


Fig. 10. Spectrum of HC15NO in the range 2400-3500 cm<sup>-1</sup>.

other isotopomers this particular resonance system is only weak. Thus, a change from one isotopomer to another gives us a "tuning" of every resonance.

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